



## Geomagnetic Observatory Gan

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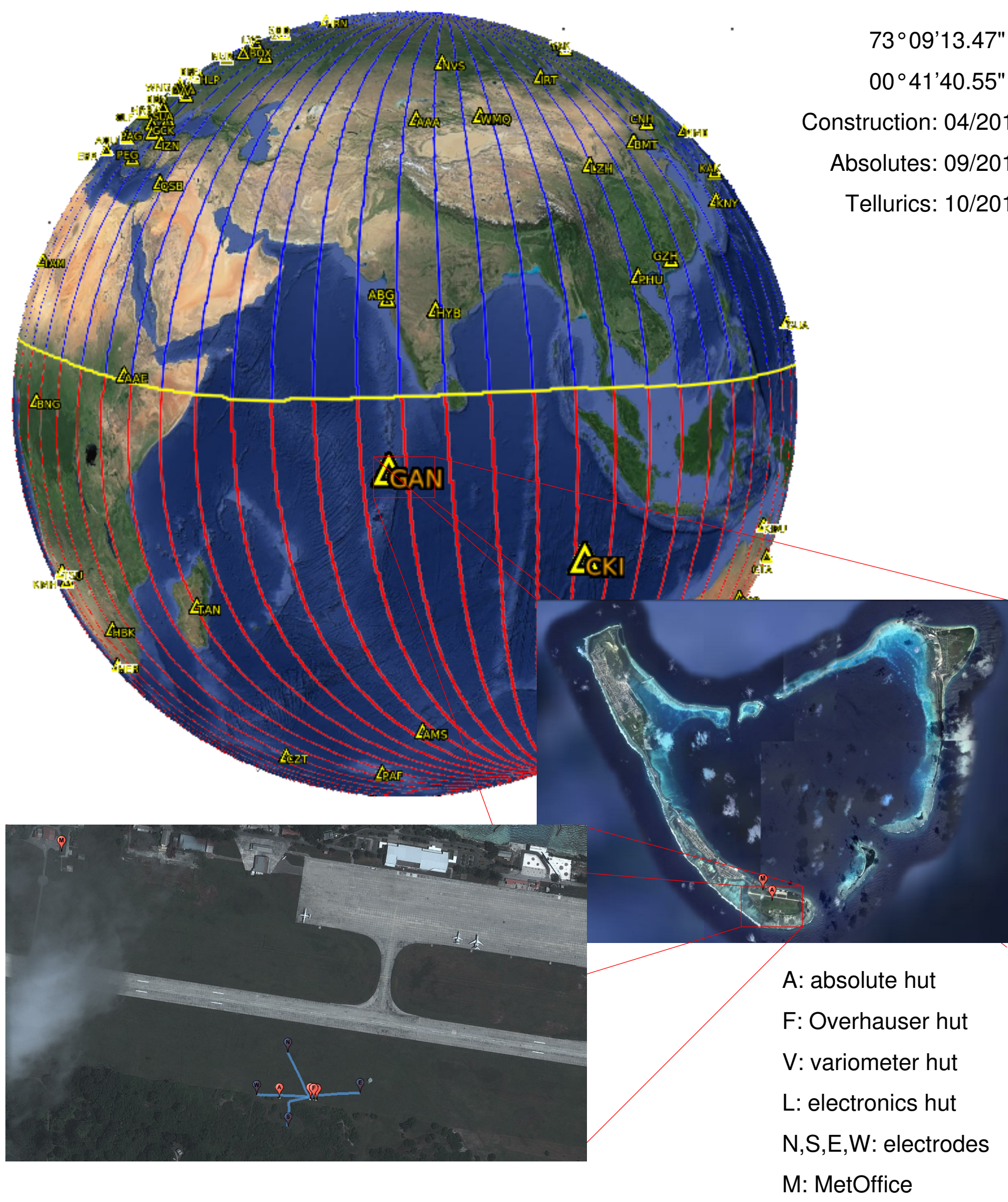




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3) National Geophysical Research Institute, Hyderabad, India, 4) Dept. of Geophysics, Charles University in Prague, Czech Republic, 5) National Space Institute, Technical University of Denmark

## Observatory location



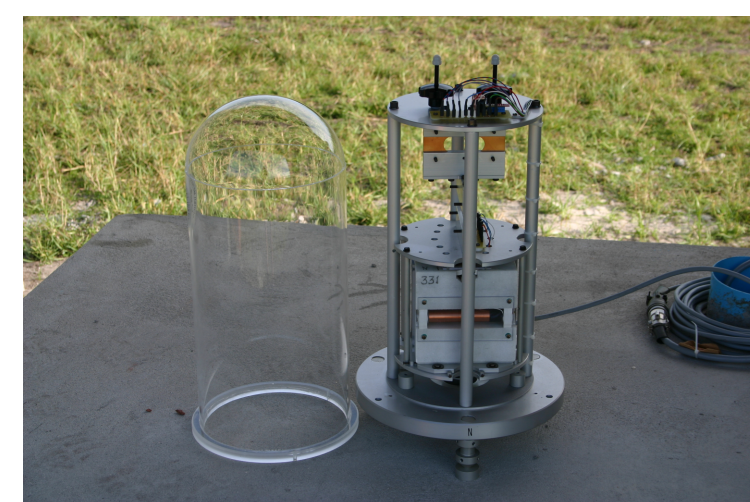
## Construction and instruments



concrete pillars 50 cm above ground level  
semi-open huts with E-W concrete block and glassfiber roofs  
glassfiber covers with styrofoam insulation, temperature and humidity monitoring  
underground plastic cable ducts  
autonomous power supply: 130 W solar panel, lead-acid batteries, voltage monitoring  
mini PC on site, WiFi link to MetOffice

MINGEO 020 declinometer/inclinometer  
(Zeiss THEO 020 + fluxgate DMI Model G FGM)suspended DMI FGM-FGE fluxgate variometer  
HDZ orientation

Gemsys GSM-90F1 Overhauser magnetometer

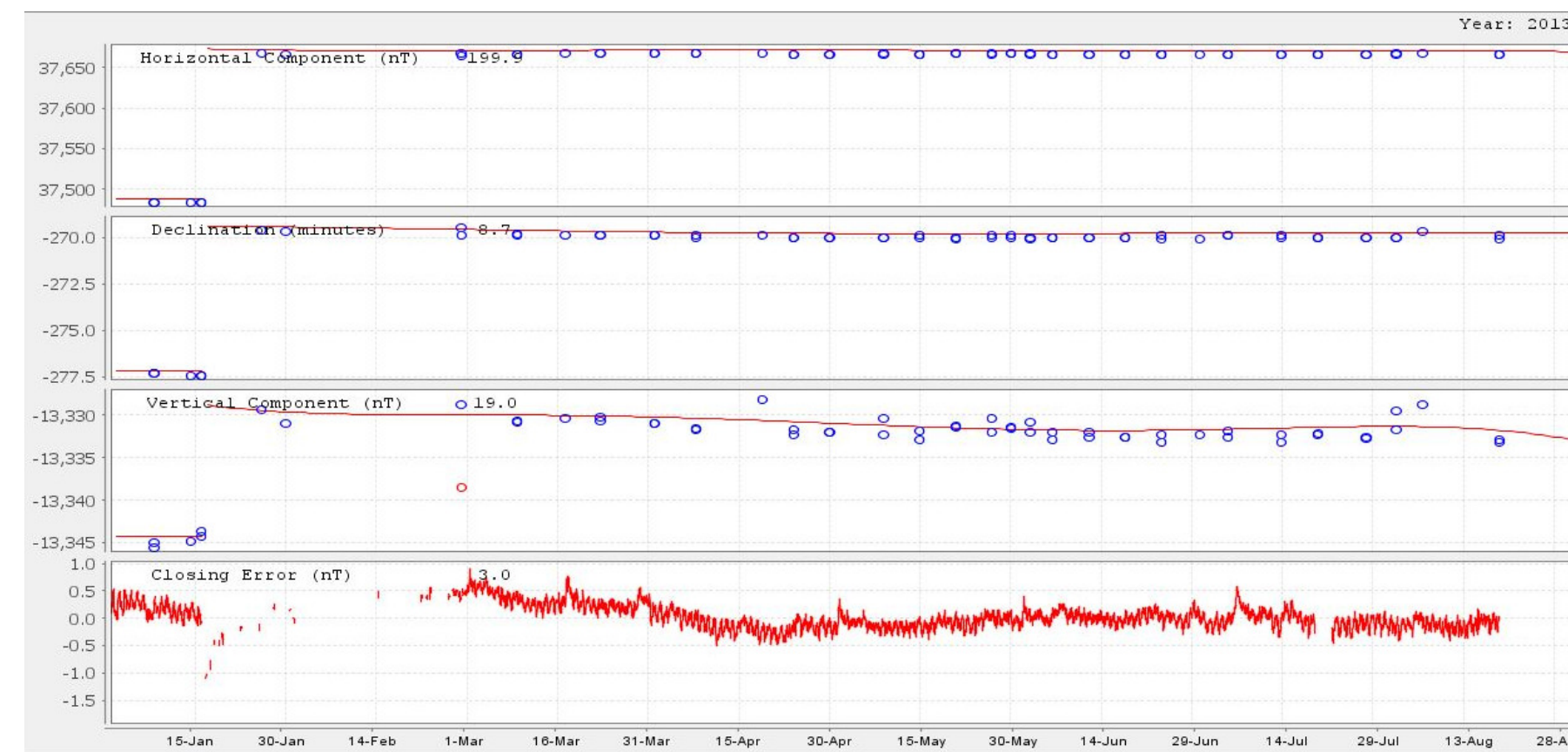
LEMI417E telluric instrument  
Pb-PbCl<sub>2</sub> electrodes, Oregon State University  
electrode spacing: N-S 146 m, E-W 200 m

Fluxgate, Overhauser, and telluric signals sampled and recorded at 1 s rate  
Minute mean data computed on site by standard Intermagnet (90 second Gaussian) filter  
Data transmitted every 10 minutes using WiFi link to MetOffice  
Local backups at MetOffice and ETH servers  
After screening and peak removal, data sent to Edinburgh GIN  
Absolute measurements performed weekly  
Observatory homepage: <http://kobizek.ethz.ch/gan.html>  
Quicklook at GAN, CKI, API magnetograms: <http://kobizek.ethz.ch/show/>

## Future plans

Sustained operations at least for the duration of the Swarm mission and hopefully beyond  
Gradual transfer of data processing knowledge and responsibilities to local staff  
Additional electrodes of different type and positioning for telluric measurements  
IntermagNet application pending

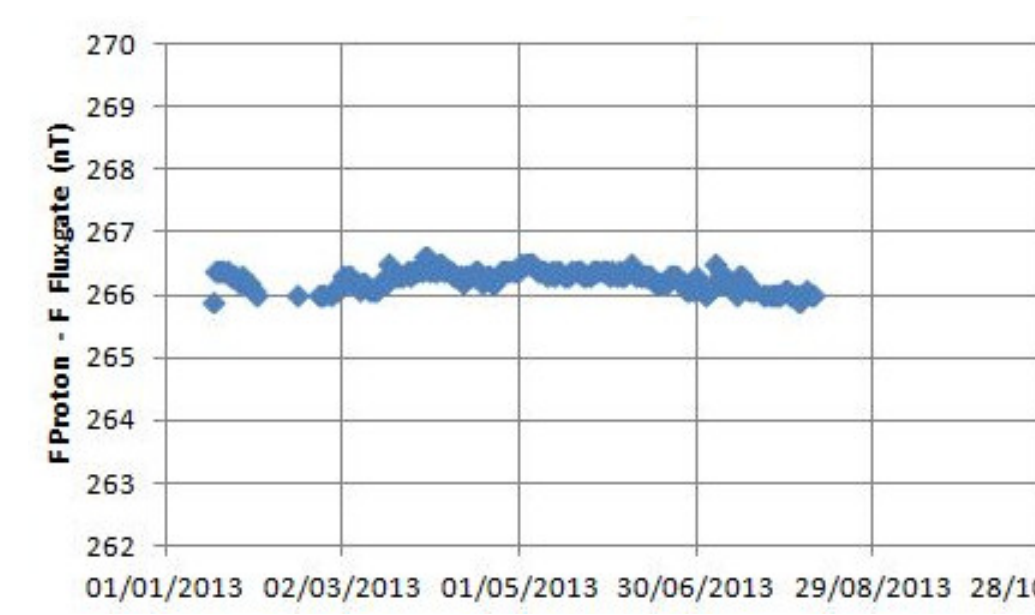
## Baseline stability



Baselines of H, D, and Z components for 2013 have stabilized after replacement of the variometer on January 15, 2013

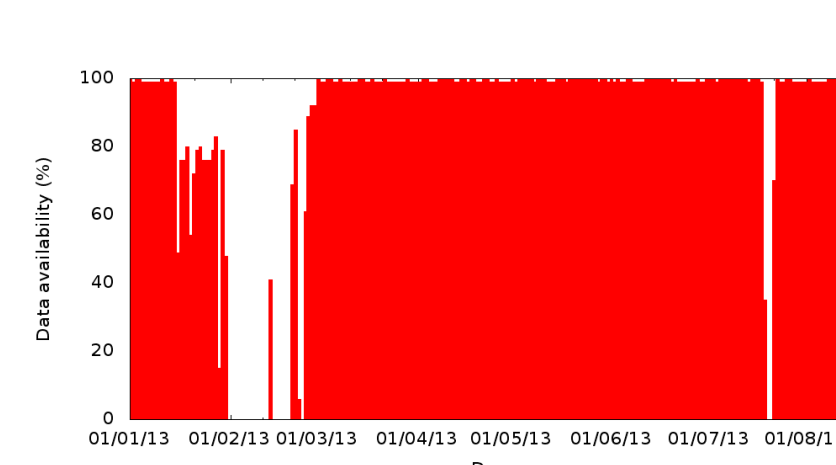
The closing error (difference between Overhauser measurement and baseline corrected fluxgate data) is also stable (shown in red)

## Stability of absolute measurements



Difference between Overhauser data and amplitude obtained from the absolute measurements indicates good absolute observations and absence of external interference due to movements of magnetic objects near the site

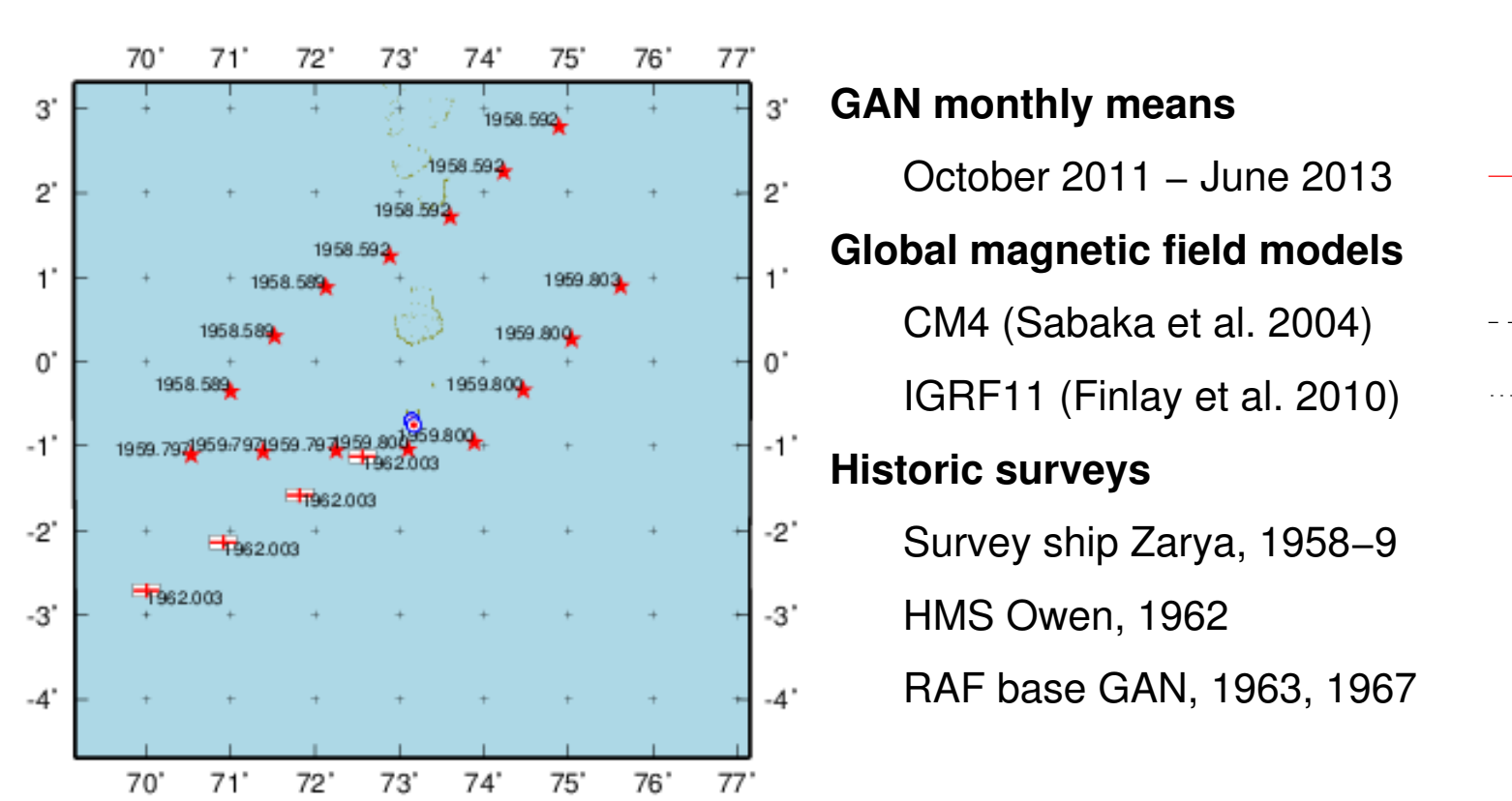
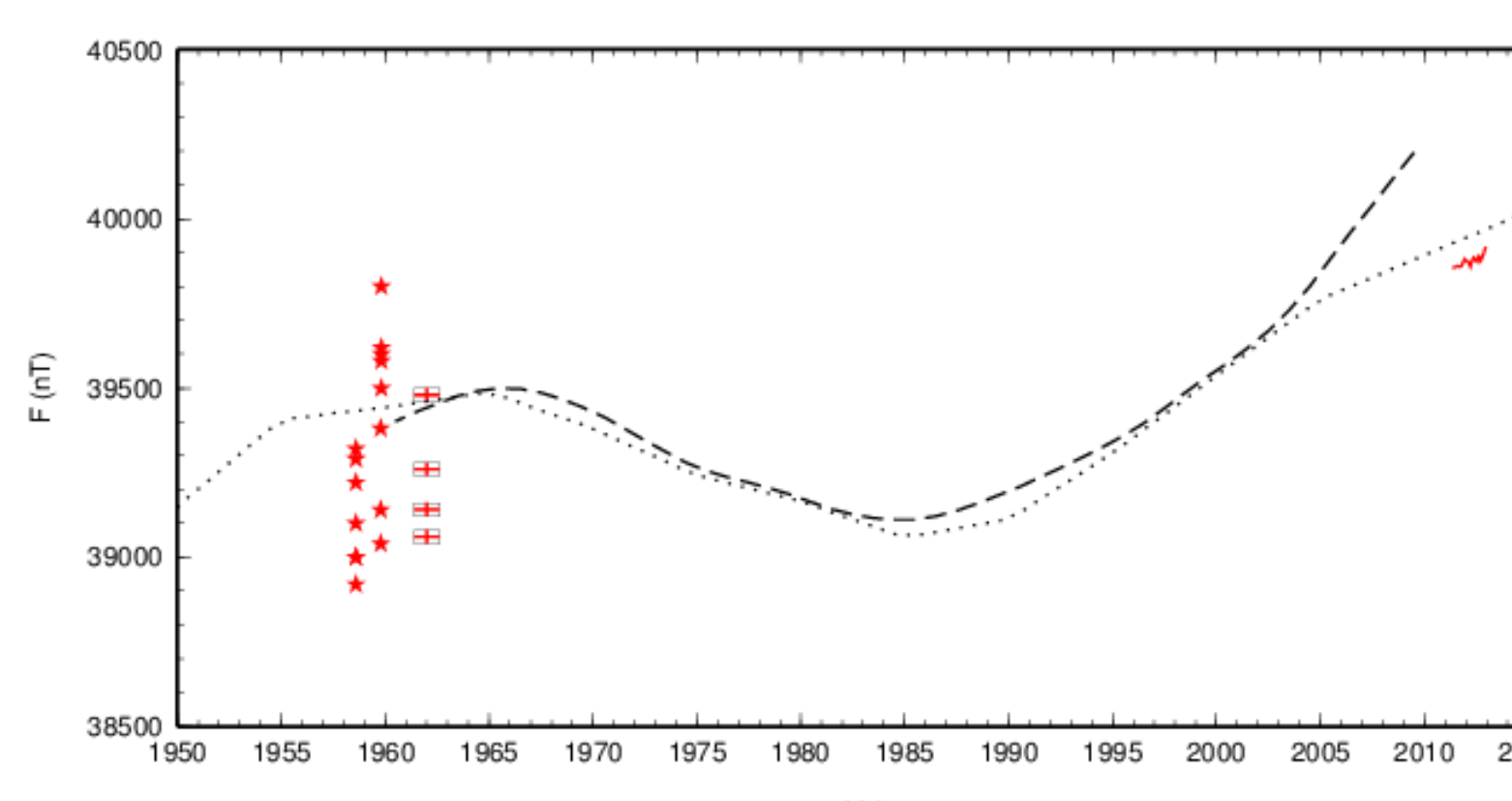
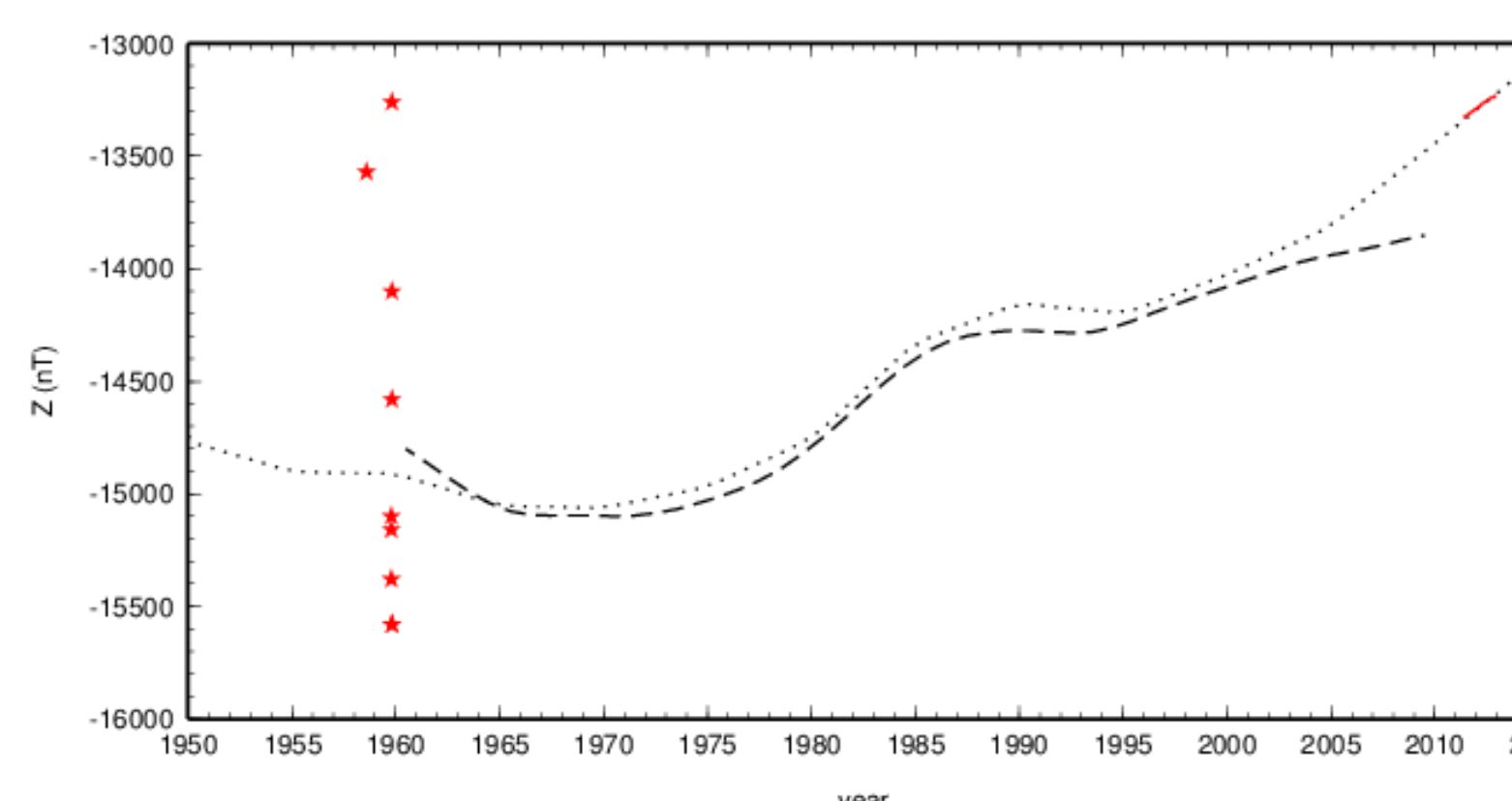
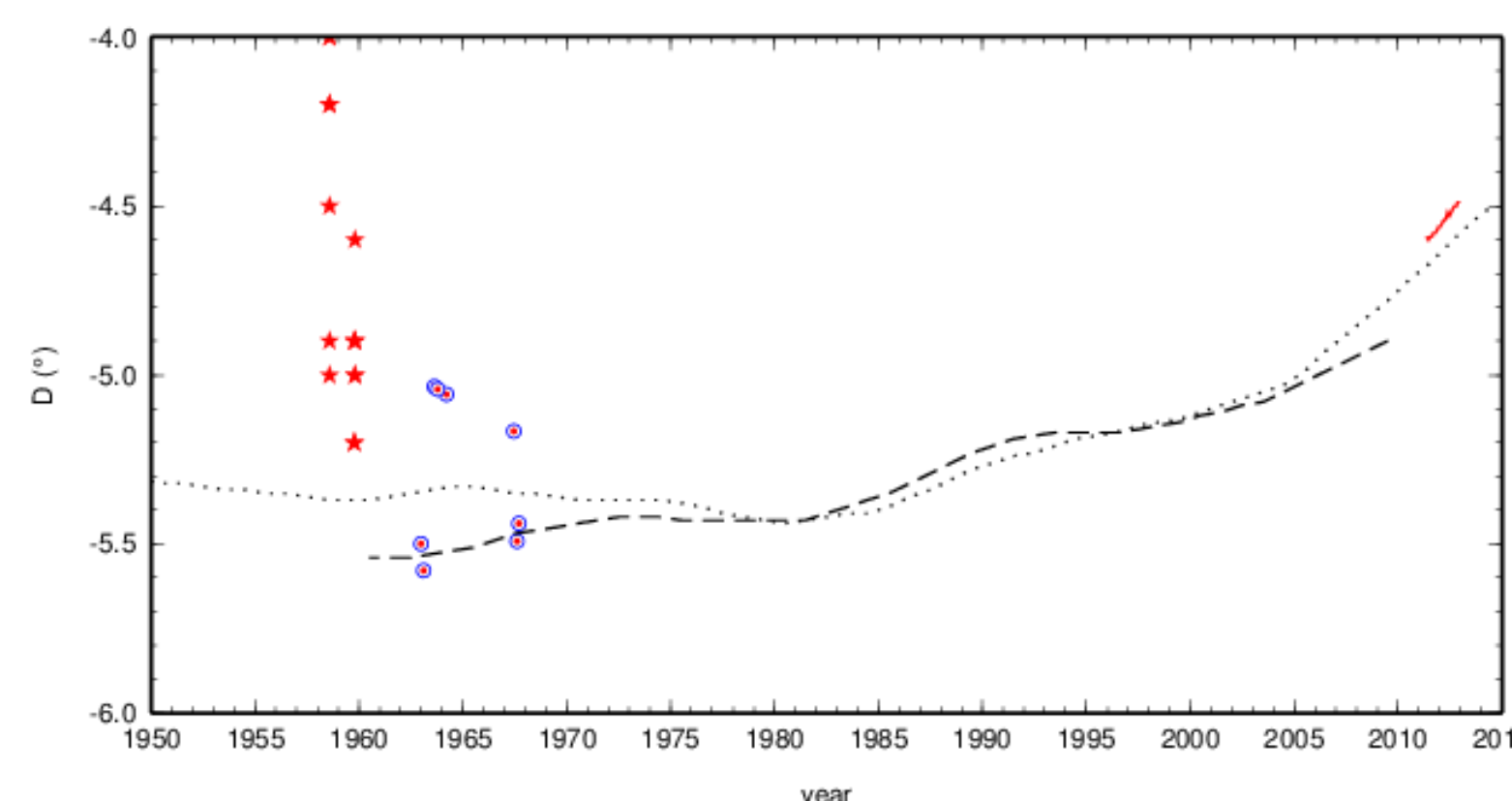
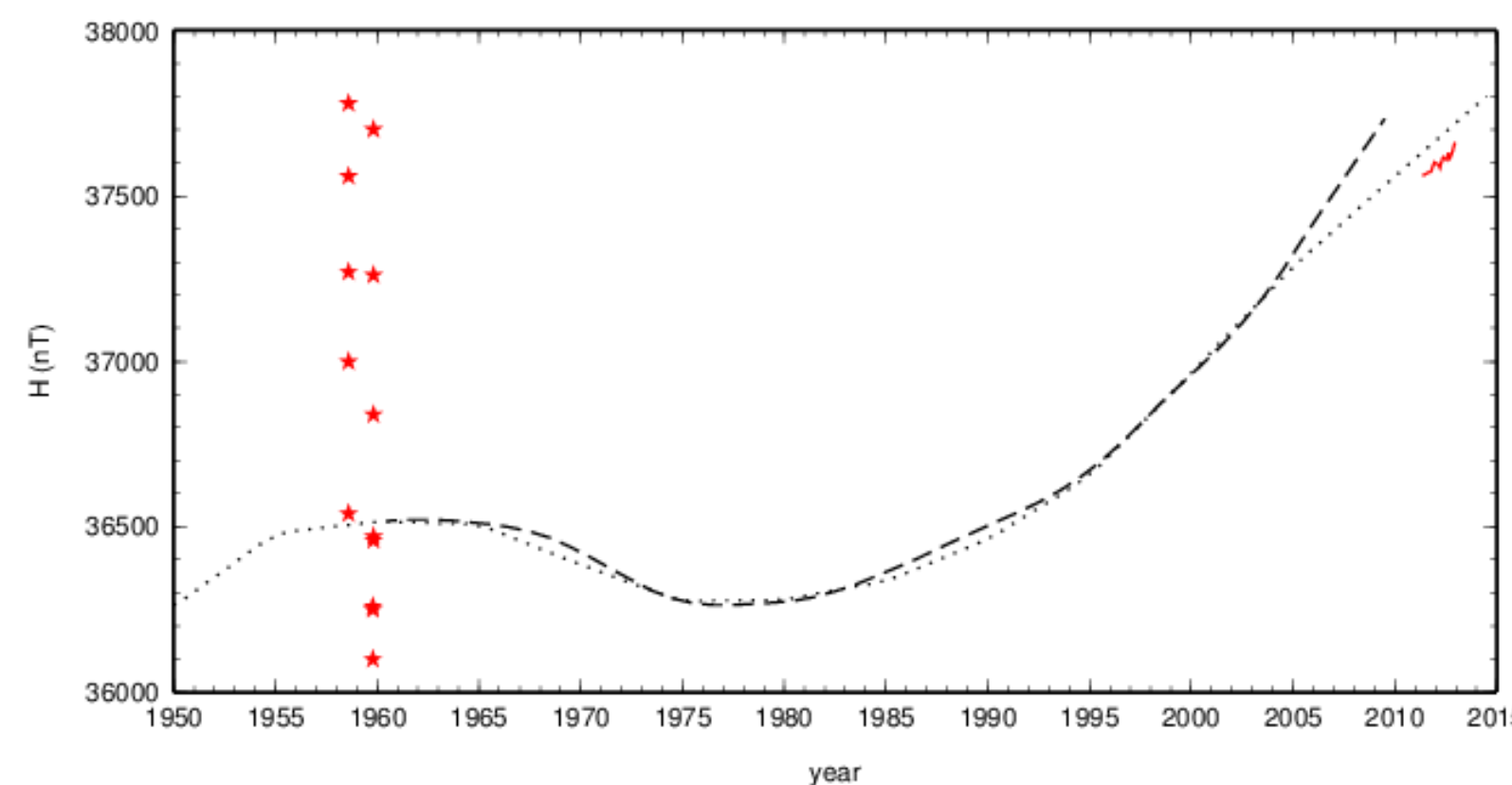
## Data availability in 2013



## Various problems and solutions

Interference between solar charger and GSM90:  
DC-DC filter installed in 2011  
Temperature stability vs. humidity in the V hut:  
Large cover with a single styrofoam layer  
Variometer drift and unstable baselines:  
New variometer installed in 01/13  
Regular grass cutting at the airport  
Peak removal from data; repair of cut cables

## Secular variation



## GAN monthly means

October 2011 – June 2013

## Global magnetic field models

CM4 (Sabaka et al. 2004)

IGRF11 (Finlay et al. 2010)

## Historic surveys

Survey ship Zarya, 1958–9

HMS Owen, 1962

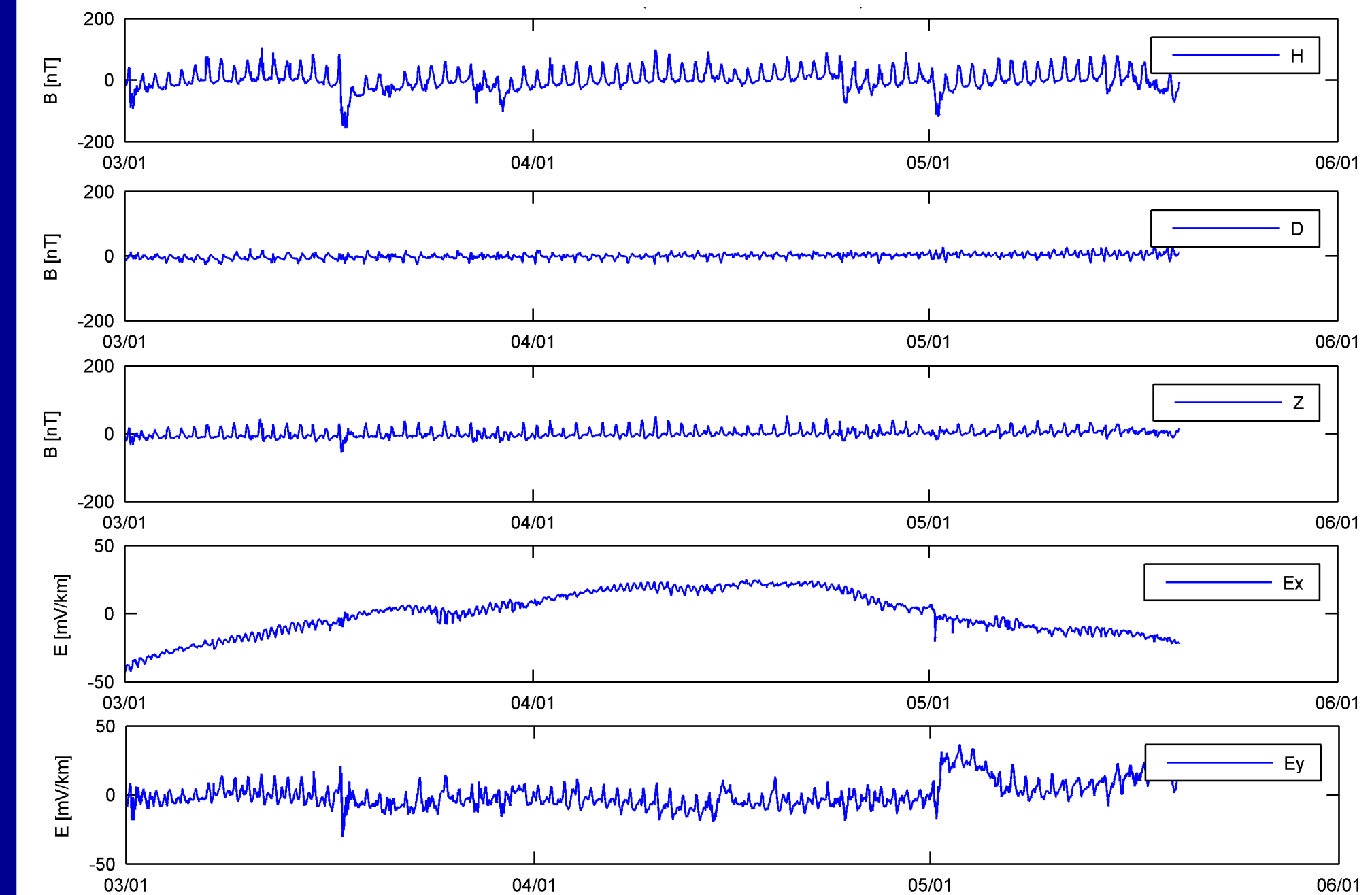
RAF base GAN, 1963, 1967

## Concluding Remarks

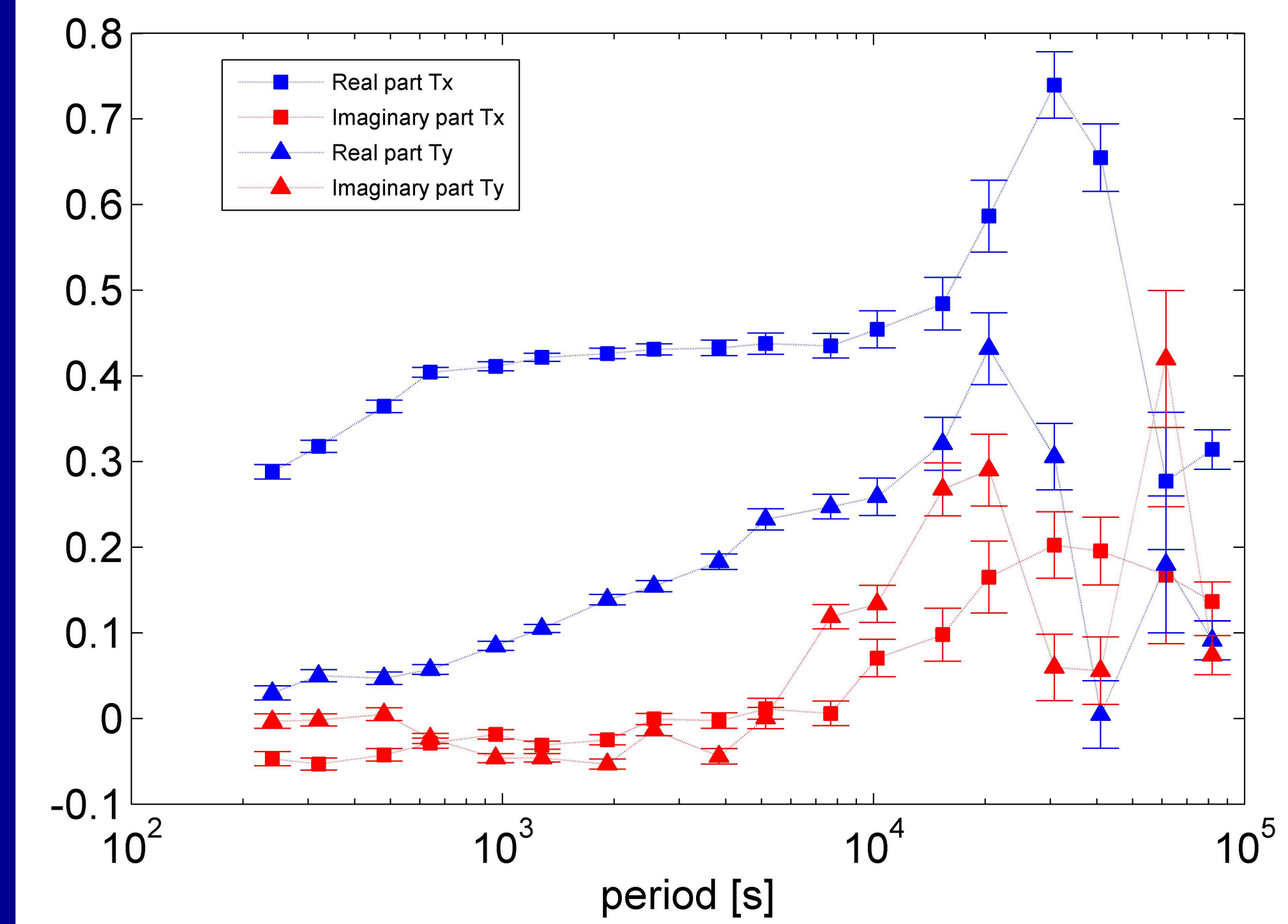
Differences between IGRF11 and actual data up to 100 nT in H, and 0.1° in D

## Preliminary EM induction analysis

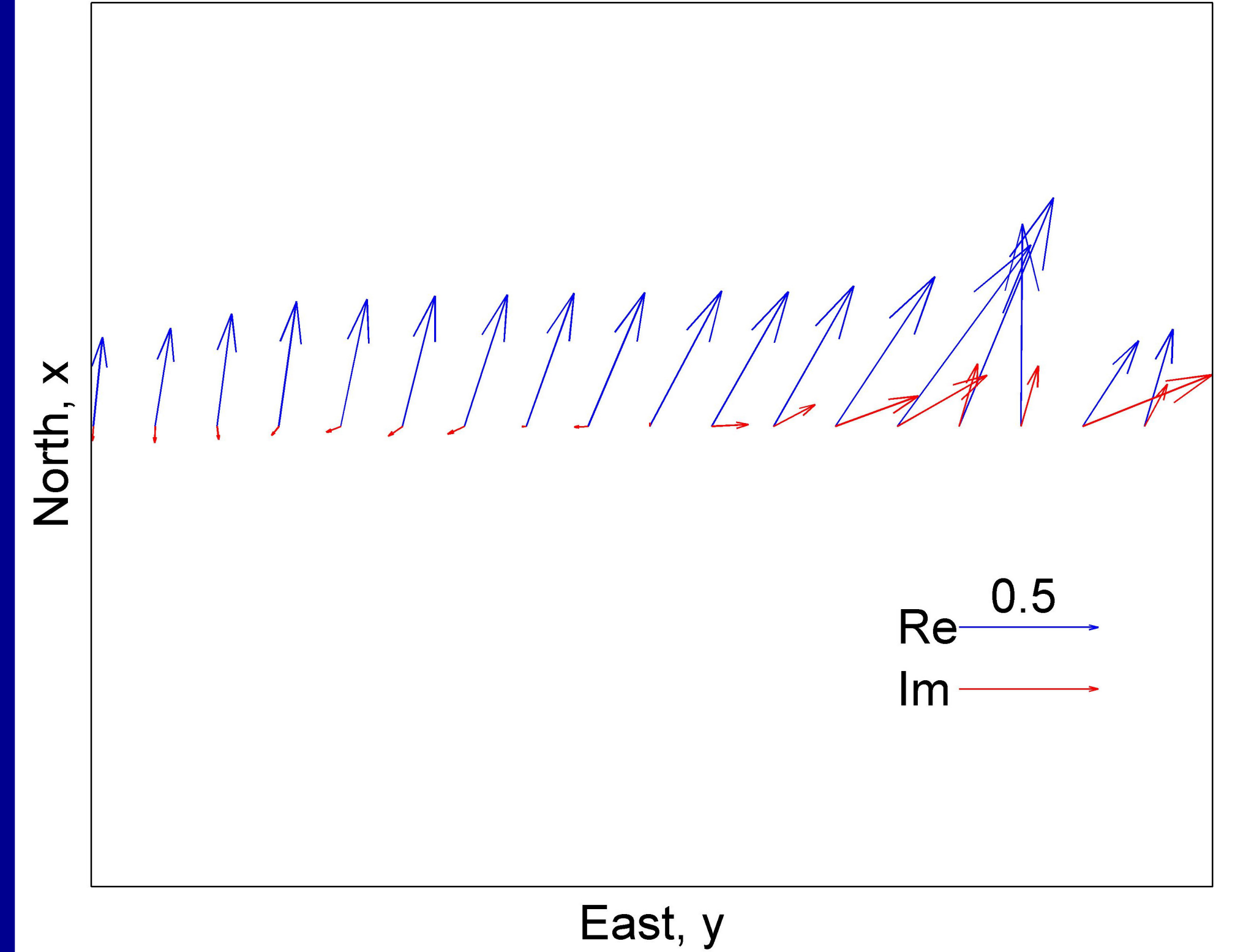
## Minute means of magnetic and horizontal electric fields, March–June 2013



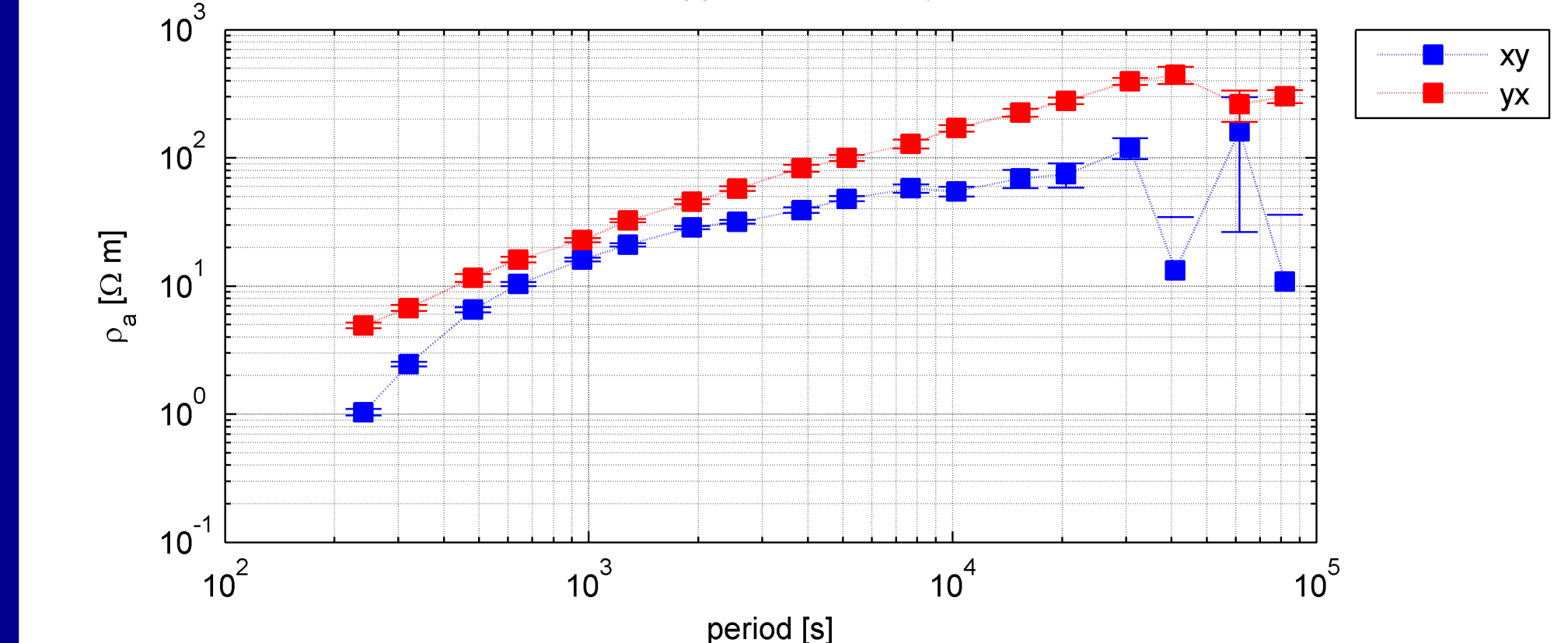
## Tippers



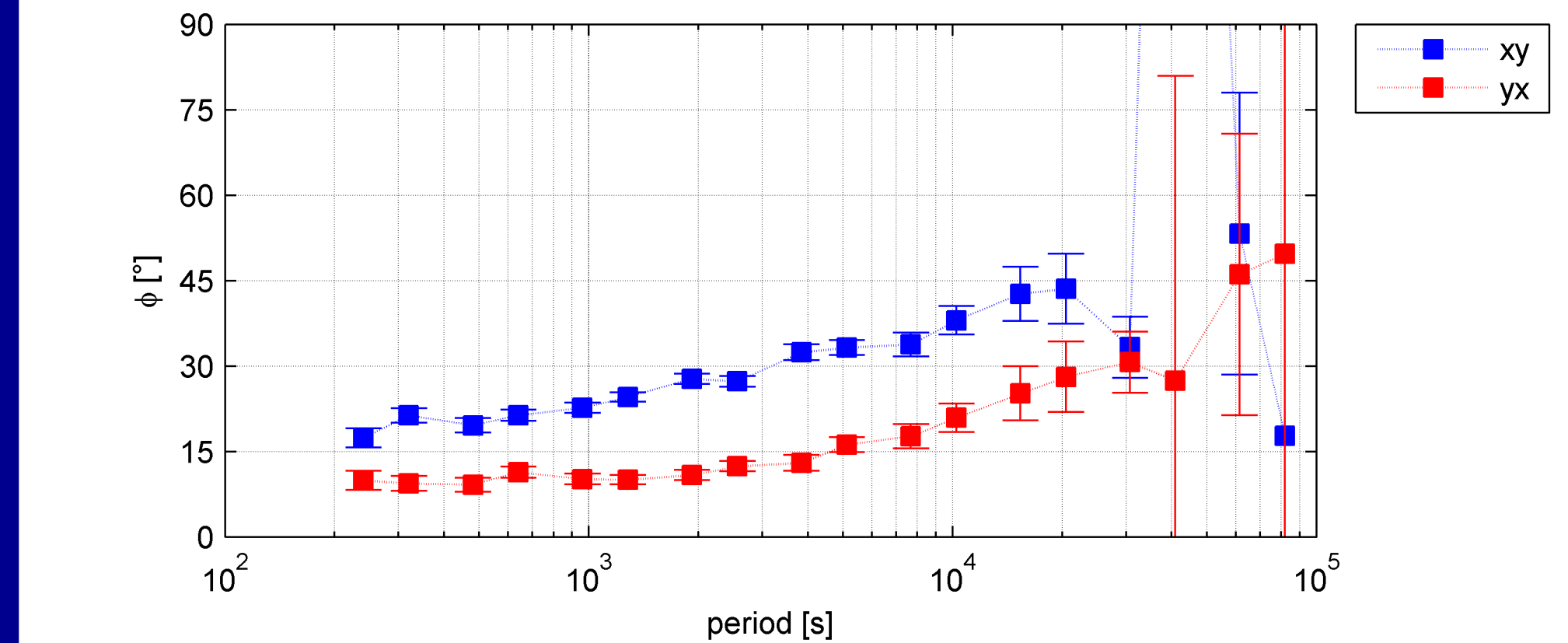
## Tippers (arrow representation)



## Apparent resistivity



## Phase



## Concluding Remarks

Minute means processed by BIRRP (Chave &amp; Thomson, 2003)

Stable B-field data

Ex component is less stable, has smaller dynamic range and larger drift than Ey

Tippers are stable for periods up to 3 hours, distorted by source effect for longer periods

Tippers point away from the deep conductive ocean

Apparent resistivities are stable for periods up to 8 hours,

distorted by source effect for longer periods

Lower stability and signal to noise ratio of Ex component

increase the scatter of xy apparent resistivity